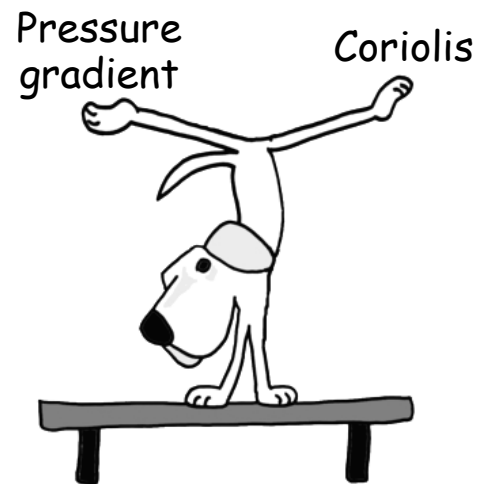


Forces

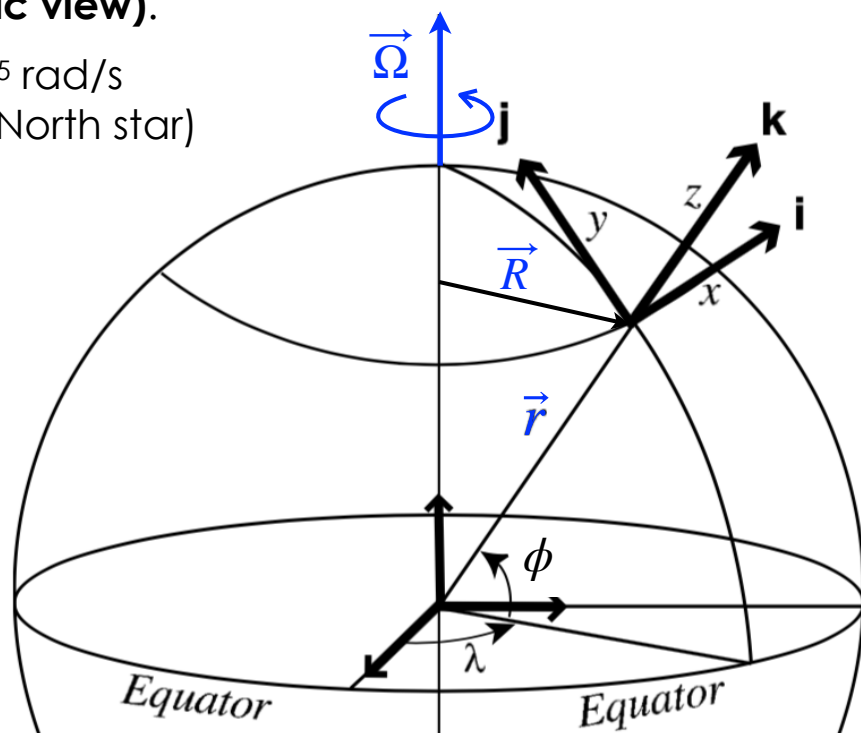
1. Rotating frame
2. Coriolis force
3. Centrifugal force



Rotating Frame

The earth rotates with a period about 24 hours (23h 56m; or 366 rotation per year), **although we can not feel it (human love geocentric view).**

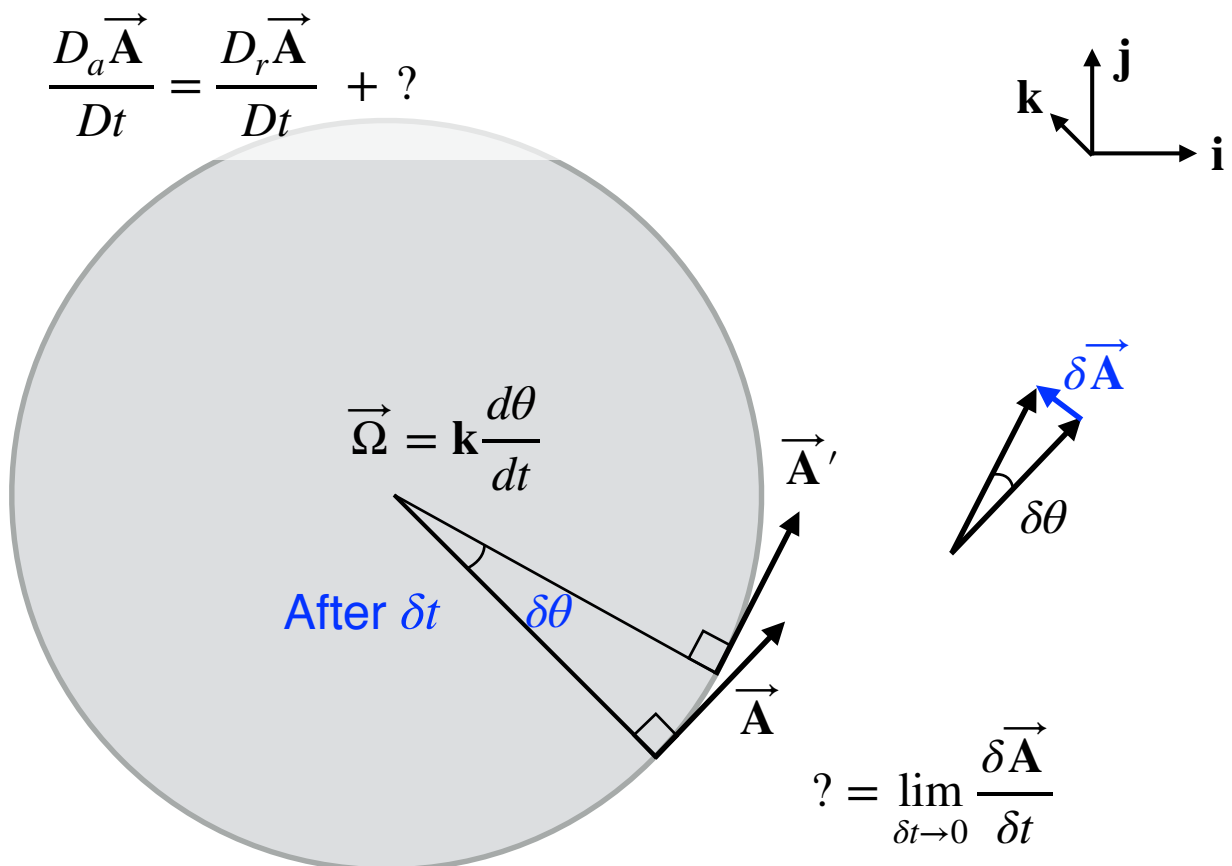
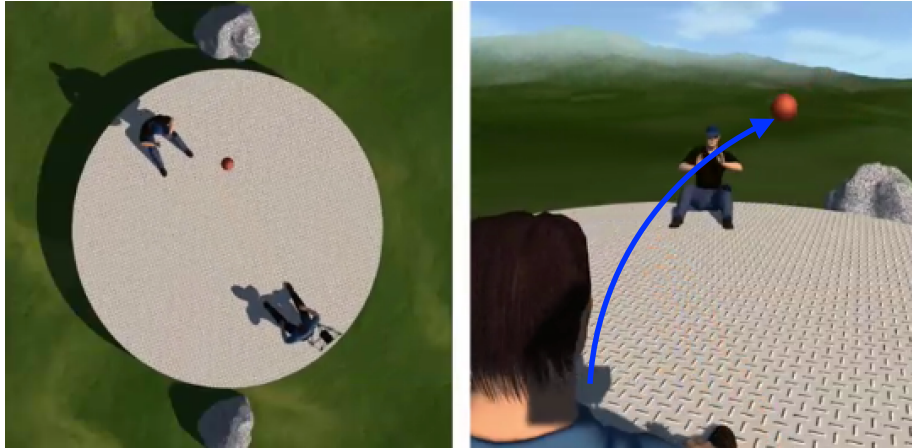
$\Omega = 7.292 \times 10^{-5} \text{ rad/s}$
(directing the North star)

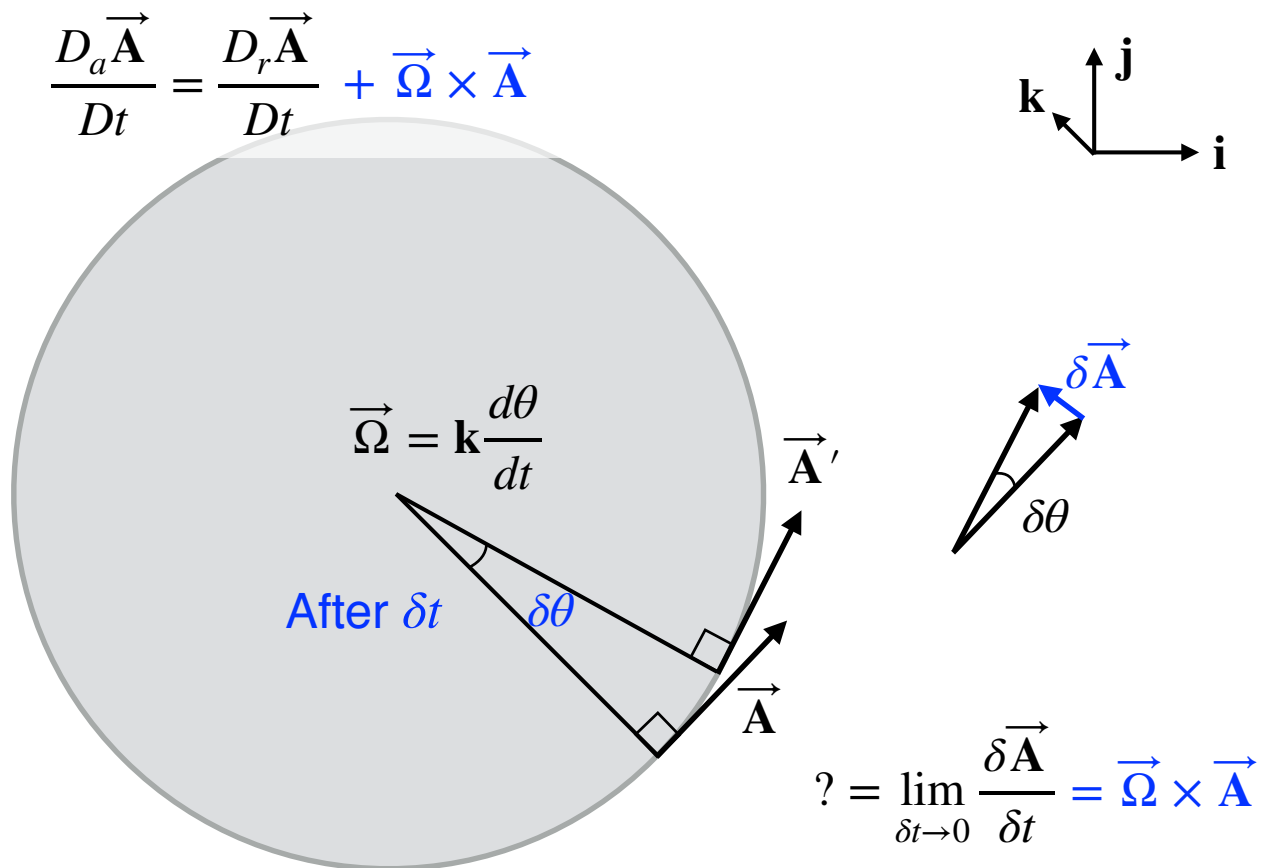


Coriolis force

Newtons' laws of motion work only in an inertial frame. However, we can still use Newton's law in a rotating frame in a uniform rotation, if we introduce two apparent forces:

1) Coriolis and 2) Centrifugal forces





Inertial frame vs. rotation frame

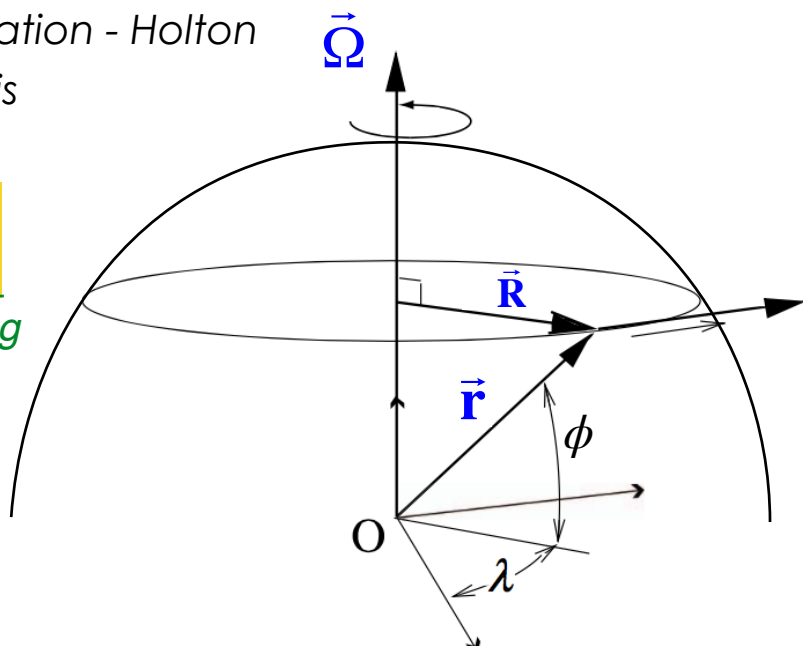
Relationship between total differentiations in a inertial frame ($D_a \vec{A}/Dt$) and a rotating frame ($D_r \vec{A}/Dt$)

Can be obtained through

- (1) vector calculation - Holton
- (2) intuition - Vallis

$$\frac{D_a \vec{A}}{Dt} = \frac{D_r \vec{A}}{Dt} + \vec{\Omega} \times \vec{A}$$

relationship working
for any vector \vec{A}



Velocity (wind)

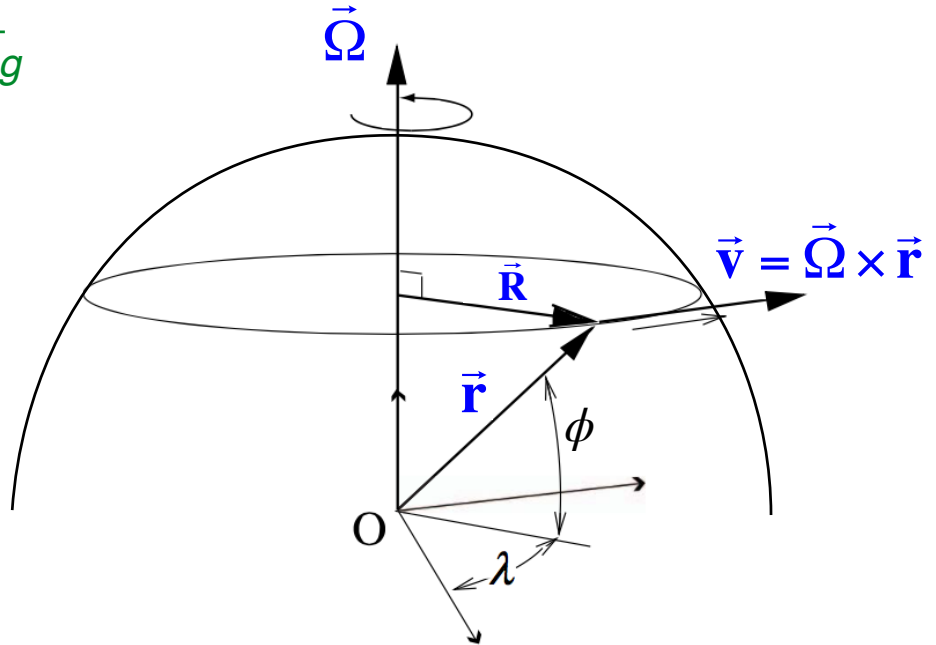
Relationship between total differentiations in
a inertial frame ($D_a \mathbf{A}/Dt$) and a rotating frame ($D_r \mathbf{A}/Dt$)

$$\frac{D_a \vec{\mathbf{A}}}{Dt} = \frac{D_r \vec{\mathbf{A}}}{Dt} + \vec{\Omega} \times \vec{\mathbf{A}}$$

*relationship working
for any vector \mathbf{A}*

$$\frac{D_a \vec{\mathbf{r}}}{Dt} = \frac{D_r \vec{\mathbf{r}}}{Dt} + \vec{\Omega} \times \vec{\mathbf{r}}$$

$$\vec{\mathbf{v}}_a = \vec{\mathbf{v}}_r + \vec{\Omega} \times \vec{\mathbf{r}}$$



Acceleration

Relationship between total differentiations in
a inertial frame ($D_a \mathbf{A}/Dt$) and a rotating frame ($D_r \mathbf{A}/Dt$)

$$\frac{D_a \vec{\mathbf{A}}}{Dt} = \frac{D_r \vec{\mathbf{A}}}{Dt} + \vec{\Omega} \times \vec{\mathbf{A}}$$

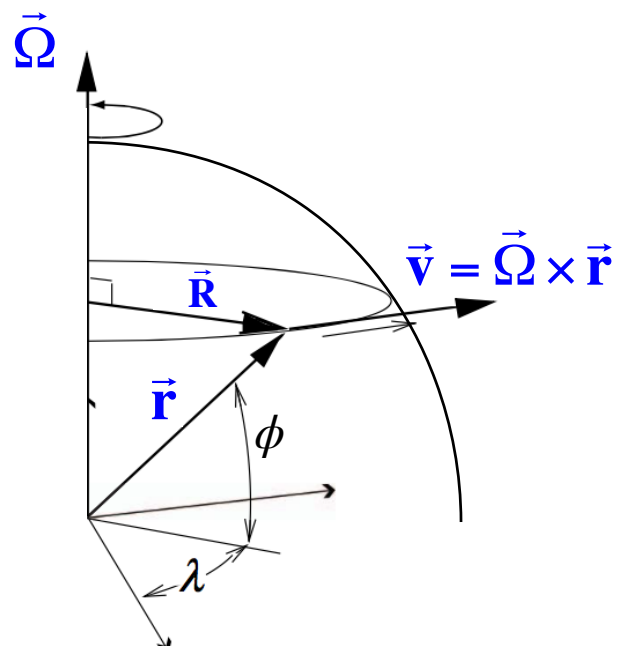
*relationship working
for any vector \mathbf{A}*

$$\frac{D_a \vec{\mathbf{v}}_a}{Dt} = \frac{D_r \vec{\mathbf{v}}_a}{Dt} + \vec{\Omega} \times \vec{\mathbf{v}}_a$$

$$\text{recall } \vec{\mathbf{v}}_a = \vec{\mathbf{v}}_r + \vec{\Omega} \times \vec{\mathbf{r}}$$

$$\frac{D_a \vec{\mathbf{v}}_a}{Dt} = \frac{D_r (\vec{\mathbf{v}}_r + \vec{\Omega} \times \vec{\mathbf{r}})}{Dt} + \vec{\Omega} \times (\vec{\mathbf{v}}_r + \vec{\Omega} \times \vec{\mathbf{r}})$$

$$\frac{D_a \vec{\mathbf{v}}_a}{Dt} = \frac{D_r \vec{\mathbf{v}}_r}{Dt} + 2\vec{\Omega} \times \vec{\mathbf{v}}_r + \vec{\Omega} \times (\vec{\Omega} \times \vec{\mathbf{r}})$$



Acceleration

$$\frac{D_a \vec{v}_a}{Dt} = \frac{D_r \vec{v}_r}{Dt} + \underbrace{2\vec{\Omega} \times \vec{v}_r}_{\text{Coriolis force}} + \underbrace{\vec{\Omega} \times (\vec{\Omega} \times \vec{r})}_{\text{Centrifugal force } (= -\Omega^2 \vec{R})}$$

$$\frac{D_a \vec{v}_a}{Dt} = -\frac{1}{\rho} \nabla p - \mathbf{k}g^* + \nu \nabla^2 \vec{v}_a$$

Handle as one term using the concept of "geoid"
(max of $\Omega^2 R$ is $\sim 0.034 \text{ m/s}^2$;
just slightly modify $g^* \rightarrow g$)

$$\frac{D_r \vec{v}_r}{Dt} = -2\vec{\Omega} \times \vec{v}_r + \underbrace{\Omega^2 \vec{R}}_{\text{viscous force is local}} - \mathbf{k}g^* - \frac{1}{\rho} \nabla p + \nu \nabla^2 \vec{v}_a$$

viscous force is local
cannot feel rotation $v_a \rightarrow v$

$$\frac{D\vec{v}}{Dt} = -2\vec{\Omega} \times \vec{v} - \frac{1}{\rho} \nabla p - \mathbf{k}g + \nu \nabla^2 \vec{v}$$

Momentum equation in a rotating frame
(with rotation rate Ω)

Equation set

Six equations with six known (u, v, w, T, p, ρ)

Mathematically closed; and **solvable**

num.
eq.

| | | |
|---|--|----------------------------------|
| 3 | $\frac{D\vec{v}}{Dt} = -\frac{1}{\rho} \nabla p - 2\vec{\Omega} \times \vec{v} - \mathbf{k}g + \nu \nabla^2 \vec{v}$ | (momentume eq. for u, v, w) |
| 1 | $c_p \frac{DT}{Dt} - \frac{1}{\rho} \frac{Dp}{Dt} = J$ | (thermodynamic energy eq. for T) |
| 1 | $\frac{D\rho}{Dt} + \rho \nabla \cdot \vec{v} = 0$ | (continuity eq. for ρ) |
| 1 | $p = \rho RT$ | (equation of state for p) |